

1 Mesoscale and high-frequency variability of macroscopic particles (> 100  $\mu\text{m}$ ) in the Ross Sea  
2 and its relevance for late-season particulate carbon export

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8

9 Abstract:

10 The Ross Sea plays a major role in the transfer of organic carbon from the surface into the deep  
11 sea due to high productivity in that system. Here we present the first particle inventory (>100  
12  $\mu\text{m}$ ) of the Ross Sea based on a combined deployment of a video particle profiler and a high-  
13 resolution digital holographic microscope. Long-distance (100s of kilometers) and short-distance  
14 (10s of kilometers) sections showed high variability of particle distributions that were dependent  
15 on the density structure of the water column. Particle export was especially apparent at sites of  
16 locally weakened pycnoclines. Similarly, *Phaeocystis antarctica* colonies that were initially  
17 retained in the mixed layer by a strong density gradient sank below the euphotic zone after  
18 erosion of the main pycnocline. Fine scale analysis at a resolution <1m revealed a significantly  
19 overdispersed (i.e., patchy) environment in all casts. Patchiness, as determined by the Lloyd  
20 index of patchiness and the Index of Aggregation, increased below the pycnocline presumably  
21 due to aggregation of particles while accumulating on density gradients. In contrast, particles in  
22 the upper mixed layer and in the nepheloid layers were more randomly distributed. In  
23 approximately half of the 84 video depth profiles, a periodicity of particle peaks was detectable,  
24 ranging from 10 to 90 m with a mode of 30 m, which can be regarded as the “relevant scale” or  
25 “characteristic patch size” of the vertical distribution of particles. We speculate that this banding  
26 of particles are sinking events that reflect cyclical weather patterns and changing wind speeds.

27 While chlorophyll fluorescence and particle mass determined by the video particle profiler were  
28 significantly correlated, the relationship changed from station to station and through time,  
29 reflecting changes in the relative contribution of fresh phytoplankton to total particle mass.  
30 Particles that sank below the main pycnocline were composed of diatoms, marine snow with and  
31 without embedded phytoplankton, crustacean plankton, and a surprisingly high percentage of  
32 heterotrophic (and perhaps mixotrophic) protists such as acantharians and tintinnids.

33

### 34 1. Introduction

35 The nutrient-rich Ross Sea is the site of massive seasonal blooms of phytoplankton (primarily  
36 diatoms and *Phaeocystis antarctica*) and accumulation of dissolved and particulate organic  
37 carbon (Carlson et al. 2000, DiTullio et al. 2000). Particle export in the Ross Sea has commonly  
38 been measured with sediment traps (e.g., Asper and Smith 1999, Smith and Dunbar 1998,  
39 Accornero et al. 1999). The POC inventory of the water column has been measured on several  
40 expeditions using Niskin bottles and GF/F filters that capture all particulate matter  $> \sim 0.7 \mu\text{m}$   
41 (e.g., Carlson et al. 2000). However, particles between  $50 \mu\text{m}$  to several millimeters contribute  
42 most to the mass flux as smaller particles do not sink sufficiently fast, and larger particles are too  
43 rare to play a major role (Guidi et al. 2008, McDonnell and Buesseler 2010). The smallest  
44 particles in this size range are primarily composed of single diatom cells ballasted by their silica  
45 skeletons (McDonnell and Buesseler 2010). Surveys of these critical larger particles are much  
46 rarer, especially in the Ross Sea (Asper and Smith 2003). Since optical backscatter and beam  
47 transmissometry are more responsive to fine particles and colloidal material (Battisto et al. 1999,  
48 Bochdanksy et al. 2010), large particles are most efficiently measured by camera systems  
49 (Stemmann et al. 2000, Guidi et al. 2008, Iversen et al. 2010). In order to better understand the  
50 spatial and temporal distribution of macroscopic particles in the late season Ross Sea, we  
51 deployed a video particle profiler (VPP) in combination with a digital holographic microscope  
52 (DIHM).

53

54

## 55 2. Methods

### 56 2.1 Research expedition details

57 Data presented in this manuscript were collected on the RVIB *Nathaniel B Palmer* from  
58 February 12 to March 16, 2013. The main focus was on the western Ross Sea as it represents a  
59 significant site for Antarctic Bottom Water formation. During that time period, the Ross Sea  
60 transitioned from being almost entirely ice free to almost completely ice covered. We focused on  
61 three areas in the western Ross Sea: north of Franklin Island, South of Coulman Island, and Terra  
62 Nova Bay, each of which we revisited several times during the expedition in order to record  
63 temporal changes. We also sampled along 76° 30'S across the width of the Ross Sea.

### 64 2.2 Video particle profiler (VPP)

65 The VPP is similar to that published in Bochdansky et al. 2010. However, instead of lighting at  
66 45° angle from both sides, side lighting with two white high-intensity LED lights was used ~7  
67 cm in front of the lens. The light beams were restricted using a slit of 1 cm in width; however, as  
68 the light intensity dropped exponentially in the front and the back of the image beam, only the  
69 brightest lit image plane was used for analysis. This method reduced bias due to overlapping  
70 particles and provided more accurate particle size estimates. At the focal plane, the imaged area  
71 was 3 cm tall and 4 cm wide. The analysis program for the video particle profiler was expanded  
72 from that in Bochdansky et al. (2010) to include more variables for particle characterization  
73 (including perimeter, volume and porosity). The VPP can record 30 images per second, with  
74 image analysis by a Linux-based image analysis program (a converted Avidemux video editor) at  
75 high speeds (approximately in real time). The images were later aligned with depth from the  
76 CTD using time as the common variable and by filming a clock displaying UTC at the beginning  
77 and the end of each video sequence. In Matlab, CTD data were matched line-by-line with each  
78 set of the particle data. The raw data consisted of millions of particles each with its own set of  
79 associated CTD data. These raw data allow us to resample particle metrics at all possible scales.  
80 Particle volumes were calculated as shown in Fig. 1, avoiding the bias of assigning  
81 disproportionately large volumes to elongated objects. Total particle volume per meter depth  
82 interval was approximated by multiplying the mean volume of particles with the mean particle  
83 number at meter intervals, and expressed as parts per million (ppm) of the survey volume.

## 84 2.3 Digital inline holographic microscopy (DIHM)

85 Details of this method were published in Bochdansky et al. (2013). Briefly, a laser beam is  
86 focused on a 9  $\mu\text{m}$  single-mode optical fiber that serves as a small but intense point source of  
87 light. The expanding beam intercepts particles that create interfering shadow images on the  
88 adjacent screen of a high-resolution (4.2 megapixel) CCD camera without lens. The camera was  
89 connected to an eBOX530-820-FL1.6G-RC computer (Axiomtek) with a Gb LAN cable that  
90 recorded images on a 750 GB hard disk at a frame rate of approximately 7-12 images per second.  
91 When the laser beam intercepts a structure, a portion of the image beam scatters and interferes  
92 with the light of the main beam in a predictable pattern. This raw image represents a hologram  
93 that can then be reconstructed applying the Kirchhoff–Helmholtz transform (Xu et al. 2001) and  
94 using commercially available reconstruction software (Octopus, 4-Deep Inwater Imaging,  
95 formerly Resolution Optics). Being lens-less, the advantage of this method is that anything in the  
96 image beam can be reconstructed without having to rely on the focal plane of a lens. The entirety  
97 of the image beam volume (i.e., 1.8 ml in this configuration) can be reconstructed in this fashion,  
98 and thus explores orders of magnitude more volume than a lens-based system at the same  
99 resolution. Reconstruction of the images and analysis (particle quantities and sizes and type)  
100 were performed manually as no reliable image reconstruction and analysis system exists for the  
101 DIHM at this time. In the future, several tens of thousands of these hand-reconstructed  
102 holograms will be used for validation and calibration for automated systems. The DIHM can  
103 detect hard structures (e.g., silica, chitin, calcium carbonate, strontium sulfate) to a resolution as  
104 small as 5  $\mu\text{m}$ , and reliably count particles of 50  $\mu\text{m}$  to ~8 mm in the image volume (Bochdansky  
105 et al. 2013).

106

107 In the VPP, particle numbers cannot be directly assigned to a defined image volume because of  
108 the diffuse border of the unconstrained light beam and the fact that different particle types have  
109 different reflectivities. In contrast, the DIHM provides very accurate and precise image volumes  
110 and number of particles within the laser image beam. For these reasons, the VPP was calibrated  
111 with the DIHM at 11 stations over 10-20 m depth ranges. The conversion factor to apply to the  
112 video images was:

113 (equation 1)  $N_v = N_{dihm} \times 0.244$  (n=11 stations, SD = 0.097),

114 where  $N_v$  = number of particles > 100  $\mu\text{m}$  per video frame, and  $N_{dihm}$  the number of particles >  
115 100  $\mu\text{m}$  per ml as determined by the DIHM. In other words, each video frame corresponds to 4.1  
116 ml sample volume. In contrast to the DIHM, the VPP provides a high throughput system that  
117 allows for much greater coverage.

118

119 The degree of overdispersion (i.e., patchiness) in the system was assessed using two indices.  
120 One, the Lloyd index of patchiness (Lloyd 1967), is domain-dependent (i.e., zero values affect  
121 the estimates); the other one, the index of aggregation, is domain-independent (Bez 2000). The  
122 Lloyd index (Lloyd 1967) was calculated as:

123 (equation 2) 
$$lp = \left[ m + \left( \frac{\sigma^2}{m} - 1 \right) \right] m^{-1},$$

124 where  $lp$  is the Lloyd index of patchiness,  $m$  the mean particle density, and  $\sigma^2$  the variance of the  
125 particle density.

126 The index of aggregation (Bez 2000) was calculated as:

127 (equation 3) 
$$ia = \sum_i z_i^2 [S \times (\sum_i z_i)^2]^{-1},$$

128 where  $ia$  is the index of aggregation,  $z_i$  the particle density, and  $S$  the sample scale (set to 1 for  
129 this analysis).

130

131 Salinity, temperature, and oxygen measurements were obtained using a SeaBird 911+  
132 conductivity, temperature, and depth (CTD) probe. Salinity samples were calibrated on discrete  
133 samples at 24°C using a Guildline 8400 Autosol four-electrode salinometer. A Seabird SBE 43  
134 polarographic oxygen sensor and a Wetlabs ECO-FL fluorometer provided data on oxygen and  
135 chlorophyll fluorescence, respectively.

136

137 All raw data from the VPP and the DIHM, including the CTD context data, were archived at  
138 BCO-DMO (<http://www.bco-dmo.org/>), cross-listed under the name of the principal investigator

139 (Bochdansky) and the NSF research cruise number (NBP13-02). [Note to reviewers: We also  
140 requested archiving all the high-resolution DIHM images of which only a very small percentage  
141 have been manually reconstructed. These images represent a unique record of late-season  
142 standing stock in the Ross Sea, and could be used by investigators for future exploration of  
143 different depth ranges or taxonomic groups. In essence, they are records of micro- and  
144 mesoplankton distribution without being exposed to the decay in formaldehyde or Lugol's and at  
145 much higher spatial resolution than historical plankton samples. Automatic image analyses may  
146 also come online soon so that these images can be processed at much higher speed. Because of  
147 the many terabytes of storage requirements, our request is currently being reviewed by BCO-  
148 DMO staff]

149

#### 150 2.4 Plankton identification using the DIHM

151 For detailed examination of particles and identification of plankton, we studied only depths at  
152 and below the pycnocline. The configuration of the DIHM was intended to maximize image  
153 volume and to target particles deeper in the water column (i.e., those that contribute to export  
154 flux). For surveys of surface plankton, a much smaller gap size would have to be chosen. The  
155 sequence was run and stopped when a larger particle was encountered (approximately >20% of  
156 the screen). In this fashion all large particles (> 100  $\mu\text{m}$ ) were captured but also smaller particles  
157 that are closer to the point source of the laser beam. These smaller particles were excluded from  
158 quantitative analysis later. The maximum length was determined using the measuring tool of the  
159 Octopus reconstruction software. This result was converted to a corrected length based on object  
160 distance from the camera. The 100  $\mu\text{m}$  threshold corresponds well with the approximate  
161 minimum particle size as seen by the video camera. *Phaeocystis* colonies, because of their dense  
162 structure, do not reconstruct very well; however, they have a very characteristic shape and  
163 texture even in the unreconstructed holograms (Fig. 2). We were therefore able to perform a  
164 detailed analysis on them on all casts through all depths, but excluded them from analysis in the  
165 more restricted depth ranges (see above).

166 Section plots were created using Ocean Data View (Schlitzer 2015).

167

### 168 3. Results and Discussion

169 The emphasis of this late-season survey was on three areas in the western Ross Sea (Terra Nova  
170 Bay, south of Coulman Island, and north of Franklin Island), which were revisited several times  
171 during the research expedition to obtain information on temporal changes of the system. Terra  
172 Nova Bay was the site of highest drawdown of CO<sub>2</sub> of all sites visited (DeJong et al., 2015). We  
173 also performed short-distance transects to obtain insight on the high-resolution spatial variability,  
174 and one long-distance zonal transect across the Ross Sea at the 76° 30'S line, a section visited  
175 during previous research cruises (e.g., Carlson et al. 2000, Smith et al. 2013)

176

#### 177 3.1 Long-distance transect at the 76° 30' line

178 Using the first derivative of potential density (referred to here as sigma theta') provided insights  
179 into the strength of the density discontinuities. Along the transect, the main pycnocline was  
180 weaker in some areas than others (i.e., the bands of sigma theta' were narrower and vertical  
181 density gradients weaker). At these sites, increased particle numbers were observed at depth and  
182 frequently appeared in bands (arrows in Fig. 3). This result was a consistent pattern along the  
183 entire 76° 30'S line, although surface salinity decreased continuously eastwards. This outcome  
184 confirms that relative density changes are more important for sinking and aggregation of  
185 particles than absolute density values. Layers of marine snow and thin layers of phytoplankton  
186 are often associated with strong pycnoclines where particle maxima can be found in or just  
187 below the most pronounced density discontinuity (MacIntyre et al. 1995, Deksheniaks et al.  
188 2001). Total particle volume (ppm) was also increased in the nepheloid layer in a broad band  
189 above the ocean bottom. Apparent oxygen utilization (AOU) increased sharply below the main  
190 pycnocline as a result of oxygen-enriched surface water not being mixed below the density  
191 gradient as well as oxygen consumption of deep particle-associated heterotrophs (Fig. 3). The  
192 highest AOUs also coincided with the distribution of the largest particles, indicating that oxygen  
193 consumption was highest at sites where marine snow was found in large quantities (Fig. 3).  
194 Patchiness as indicated by the Lloyd index was highly variable throughout the water column,  
195 with highest values generally at stations and depths at which particle abundance was low. In  
196 some spots, extremely high values coincided with the largest mean particle sizes (Fig. 3).

197

### 198 3.2 North of Franklin Island short transect

199 This transect provided some insights into more highly resolved spatial variability of the Ross Sea  
200 system. Again, high particle mass in the surface was associated with a weakened pycnocline,  
201 likely the result of mixing of nutrients to the surface and a localized bloom (Fig. 4). Increased  
202 particle mass at depth was the result of a combination of high particle load at the surface and the  
203 lessened density barrier. This finding suggests that particle export this late in the season is highly  
204 episodic in nature. A strong pycnocline in one portion of the transect kept particles suspended at  
205 and just below the density gradient (left side in Fig. 4), contributing greatly to the high AOU  
206 there (Fig. 4). Particle size increased at and just below the pycnocline likely as a result of  
207 aggregation (Alldredge et al. 2002).

208

### 209 3.3 South of Coulman Island

210 A massive particle export was apparent in bands at one station in the center of the line that  
211 coincided with a weak pycnocline (Fig. 5a, arrow). However, in one portion of the transect (to  
212 the left in Fig. 5), strong particle peaks were present at depth (Fig. 5b) despite a rather strong  
213 density gradient at the main pycnocline (Fig. 5a).

214

### 215 3.4 Fine-scale distribution

216 High-resolution and high-frequency analysis of particles by the VPP allowed us to explore  
217 submeter patchiness. These very small scales are highly relevant for plankton dynamics and  
218 biophysical processes (Wolk et al. 2004). Our analysis was based on individual frames collected  
219 at a frame rate of 30 per second, each of which surveys approximately 4 ml of seawater. This  
220 meant that the scale of our fine-scale analysis was well below 1 m. Whether or not particle  
221 distributions at this scale were significantly overdispersed was tested by comparing the  
222 frequency distribution of particle counts with a standard Poisson distribution. Particle  
223 distributions were significantly overdispersed ( $p < 0.005$ ) in all profiles at these small scales,

224 especially below the pycnocline and above the nepheloid layer. The two indices (the index of  
225 aggregation and the Lloyd index of patchiness) showed similar trends, but they were also  
226 markedly different from each other (Fig. 6). Sometimes the relationship bifurcates; the IA seems  
227 to flatten where the Lloyd index shows large deviations at higher values (Fig. 6). We therefore  
228 conclude that the Lloyd index is a better metric in describing highly overdispersed micropatches.  
229 A Lloyd index of 1 reflects a random distribution of particles as was the case in the presence of  
230 finely suspended material such as found in the upper mixed layer and in the bottom nepheloid  
231 layer (Fig. 3 and 4). When subsampling the same data over increasingly larger spatial scales (1  
232 m, 3 m, 5 m, 11 m, etc.), the Lloyd index of patchiness quickly approached values close to one,  
233 indicating a more random distribution of particles at these larger scales (not shown). Arbitrary  
234 binning into larger and larger depth intervals, however, misses the periodicity that was apparent  
235 in the particle peaks through the water column. For this reason, we tested for periodicity by using  
236 fast Fourier transforms (FFT) for each cast. Of 84 casts, 40 showed a clear peak in the  
237 periodogram after a FFT (example in Fig. 7b). Casts that did not show periodicity of smaller  
238 particle peaks often displayed one or two very prominent peaks indicative of large settling events  
239 (Fig. 7c). For casts with detectable periodicities, average peak-to-peak distances ranged from 10  
240 to 90 m with the most frequent bin of 30 m (Fig. 8). These periodicities can be interpreted as the  
241 “characteristic patch size” for particle peaks ( $> 100 \mu\text{m}$ ). Given the relevance of a continuum of  
242 temporal and spatial scales to phytoplankton growth (Harris 1980), even small observed  
243 periodicities require overwhelming physical or biological forcing mechanisms. In other words,  
244 even subtle observed frequencies may indicate very strong causes. While intriguing, it is unclear  
245 what caused the observed periodicity in particle peaks. These peaks either were the result of  
246 water column density structure (i.e., particles being retained at local pycnoclines), or of episodic  
247 sinking events, or both. Smith et al. (2010) observed high temporal variability of fluorescence in  
248 surface water that they attributed to wind-induced advective changes. It is thus possible that  
249 periodicity in weather patterns would lead to the observed periodicities in particle peaks with  
250 depth. This forcing would be in addition to factors known to control export such as surface  
251 production and grazer community composition (Smith and Dunbar 1998).

252

253 The relationship between particle data and fluorescence signal was also investigated.  
254 Fluorescence and ppm were correlated albeit at low  $r^2$  values (Fig. 9). At some stations, the ppm  
255 and fluorescence data were tightly coupled (Fig. 9a), while at other stations the variables  
256 diverged widely (Fig. 9b). A relatively tight coupling between fluorescence and total particle  
257 volume may be indicative of fresh phytoplankton dominating overall particle mass (Fig 9a). At  
258 other stations, the two signals were more decoupled, showing a flattening of the relationship  
259 especially at lower particle number ranges (i.e., the ranges that are most important to estimate  
260 fluxes below the primary pycnocline). One location that was probed at 2 week intervals (casts 55  
261 and 104) not only revealed a more variable relationship between total particle volume and  
262 fluorescence but also displayed significant differences in slopes and elevations of the regression  
263 lines over the two week interval (ANCOVA,  $n = 656$ , homogeneity of slopes:  $F = 1179$ ,  
264  $p < < 0.0001$ ; elevation:  $F = 339.6$ ,  $p < < 0.0001$ , Fig. 9b). Changes in the relationship between  
265 fluorescence and the total particle volume as detected by video images may thus be useful as  
266 indicators for the relative state of degradation of particulate matter, assuming that fresh  
267 phytoplankton material at the surface would have the highest fluorescence relative to particle  
268 volume, while particles dominated by heterotrophs, marine snow mucous matrix, and more  
269 refractory phytoplankton would show lower fluorescence-to-particle volume ratios. The observed  
270 decrease in fluorescence at a given particle volume later in the season is consistent with this  
271 trend (Fig. 9b).

272

### 273 3.5 Particle composition and plankton distributions

274 At 26 stations we examined the composition of large particles using the DIHM from the center of  
275 the main pycnocline downwards. For this analysis, only particles  $> 100 \mu\text{m}$  were considered  
276 because taxonomic interpretation of smaller particles is exceedingly difficult even with the  
277 DIHM. The particles were grouped into five categories: marine snow, phytoplankton,  
278 zooplankton, unidentified organisms, others (Fig. 10 and 11). Marine snow included amorphous  
279 aggregates, clusters, stringers, and aggregates containing organisms (Fig. 10). Phytoplankton  
280 included *Rhizosolenia* spp (30% of total phytoplankton observed), *Corethron* spp (7%),  
281 *Chaetoceros* spp (0.2%), other diatoms, and dinoflagellates (0.1%) (Fig. 10). Analysis of  
282 *Phaeocystis* was performed separately and in more detail (see below). Zooplankton included

283 acantharians (52% of total zooplankton observed), copepods (24%), tintinnids (9%), larvaceans  
284 (7%), and nauplii (5%) (Fig. 10). Unidentified organisms included particles clearly organismal  
285 but that could not be classified with certainty. “Others” included optically dense singular  
286 particles that did not classify as marine snow or organismal; some of them could have been large  
287 fecal pellets.

288

289 As expected, many of the larger particles were composed of marine snow that were either  
290 amorphous with unidentifiable content, or contained large numbers of aggregated diatoms (Fig.  
291 11). Particles below the main pycnocline were composed of a surprisingly large number of  
292 zooplankton, and because they cannot migrate back into the mixed surface layer (except for the  
293 largest copepods), they become part of the export flux. Acantharians as well as ciliates,  
294 radiolarians, and foraminifera are known for kleptoplasty (Stoecker et al.2009), therefore there is  
295 some overlap in the zooplankton and phytoplankton categories. Among ciliates, only loricate  
296 forms image well with the DIHM, which means that the total ciliate numbers would have to be  
297 ~10 times higher when accounting for the much more numerous aloricate ciliates (Assmy et al.  
298 2013 PNAS supplementary section Fig. 53). The lack of Radiolaria and Foraminifera in our  
299 samples was remarkable given the fact that they should produce good images with the DIHM.  
300 This result means that among Cercozoa (including former Radioloaria) and Retaria (including  
301 Foraminifera and Acantharia, Adl et al. 2005), Acantharia make up the bulk mass. These  
302 organisms were historically underestimated as they are lost with conventional preservation  
303 methods (Beers and Stewart 1970).

### 304 3.6 *Phaeocystis antarctica* colony distribution

305 *P.antarctica* (Haptophyta) colonies were very common at some stations in the western Ross Sea  
306 but remarkably confined to a region between 169° and 190°E (Fig. 13). This distribution  
307 corresponds well with the highest abundance levels observed in DiTullio et al. (2000) for earlier  
308 parts of the seasonal growth cycle. The abundance levels were highest at station # 15 followed  
309 by station # 121 (Fig. 13) both in terms of average numbers of *Phaeocystis* colonies and as  
310 integrated over the entirety of the mixed layer (Fig. 12). Over a period of two weeks, we  
311 observed sinking of a large number of colonies from the surface mixed layer through the

312 pycnocline into the deeper layer where *Phaeocystis* colonies can be considered to have been  
313 undergone export. This penetration of *Phaeocystis* into deep water was associated with a marked  
314 weakening of the pycnocline during the same period (Fig. 14). Thus *Phaeocystis* colonies not  
315 only contribute significantly to total export production during the main growing season in the  
316 Austral spring and summer (DiTullio et al. 2000) but also in the fall. Our observations, showing  
317 retention on strong pycnoclines followed by significant export, however, suggests that export is  
318 highly episodic, consistent with previous sediment trap observations (Smith and Dunbar 1998).

319

320

#### 321 4. Conclusions

322 Spatial and temporal variability of particles  $>100\ \mu\text{m}$  is very high in the autumnal Ross Sea at all  
323 measured scales. We observed a characteristic banding pattern in the microscale distribution of  
324 particles at characteristic scales of tens of meters. These bands likely reflect sinking events as a  
325 result of cyclical wind events that locally erode the main pycnocline and allow particles to sink  
326 into deeper strata. Analysis of DIHM images revealed a large contribution of live or moribund  
327 plankton below the pycnocline, which represent a source of undegraded carbon for deeper layers.  
328 Fluorometry cannot reliably determine particle mass through the water column, not only because  
329 of the limit of detection but also because of the fact that the relationship between total particle  
330 volume and fluorescence changes spatially and temporally. In turn, a tight coupling between total  
331 particle volume and fluorescence may indicate a large contribution of fresh phytoplankton to  
332 sinking fluxes.

333

334

335 Acknowledgements: We would like to thank our Co-PIs on this project Robert Dunbar, Jack  
336 DiTullio, Monica Orellana. Cody Garrison helped with the deployment of the DIHM and the  
337 VPP. We also thank the crew and the Polar Program scientific support team of the RV Nathaniel  
338 B Palmer. This project was funded by the National Science Foundation, Division of Polar  
339 Program grants #1142097 to ABB, and # 1142117 to DAH.

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411 Figure legends:

412

413 Fig. 1. Method for approximating particle volumes. Image analysis determines the projectional  
414 area of an irregularly shaped particle with length  $L$ . This area is used to determine the equivalent  
415 spherical diameter ( $ESD$ ) of a circle with the same area, which in turn is used to determine the  
416 volume of a sphere of the same diameter. The advantage of this method is that the area is  
417 dimensionally only once removed from volume, and some of the irregularities of objects are  
418 accounted for.

419

420 Fig. 2. Unreconstructed holographic image of a *Phaeocystis antarctica* colony. Diameter of *P.*  
421 *antarctica* colonies varied from  $\sim 70 - 140 \mu\text{m}$  (Tang et al. 2008).

422

423 Fig. 3. Long-distance transect along the  $76^\circ 30'S$  line (West to the left). Variables from top to  
424 bottom are: (a) the first derivative of sigma theta ( $\sigma\Theta'$ ) indicating the strength of the  
425 pycnocline, (b) the mean particle volume as the product of mean particle abundances and mean  
426 particle volume at each meter of depth (unit: ppm), (c) the apparent oxygen utilization (unit:  
427  $\mu\text{mol kg}^{-1}$ ), (d) the Lloyd index of patchiness calculated according to equation 2 (at a random  
428 distribution of particles Lloyd index = 1), (e) relative mean volume of particles (cubic pixel).  
429 Arrows indicate weakening of the main pycnocline concomitant with bands of exported particles  
430 at depth.

431

432 Fig. 4. Short-distance section north of Franklin Island (west to east direction). (a) the first  
433 derivative of sigma theta ( $\sigma\Theta'$ ) indicating the strength of the pycnocline, (b) the mean  
434 particle volume as the product of mean particle abundances and mean particle volume at each  
435 meter of depth (unit: ppm), (c) the apparent oxygen utilization (unit:  $\mu\text{mol kg}^{-1}$ ), (d) the Lloyd  
436 index of patchiness calculated according to equation 2 (at a random distribution of particles

437 Lloyd index = 1), (e) relative mean volume of particles (cubic pixel). Arrows indicate weakening  
438 of the main pycnocline concomitant with bands of exported particles at depth.

439

440 Fig. 5. Short-distance section south of Coulman Island (northwest to southeast direction). (a) the  
441 first derivative of sigma theta ( $\sigma\theta'$ ) indicating the strength of the pycnocline, (b) the mean  
442 particle volume as the product of mean particle abundances and mean particle volume at each  
443 meter of depth (unit: ppm), (c) the apparent oxygen utilization (unit:  $\mu\text{mol kg}^{-1}$ ), (d) the Lloyd  
444 index of patchiness calculated according to equation 2 (at a random distribution of particles  
445 Lloyd index = 1). Arrows indicate weakening of the main pycnocline concomitant with bands of  
446 exported particles at depth.

447

448 Fig. 6. Three representative examples of the relationship between the Index of Aggregation (Bez  
449 2000) and the Lloyd index of patchiness (Lloyd 1967) at submeter scales ( $\sim 3.3$  cm). A Lloyd  
450 index of 1 means that particles are randomly distributed, which most frequently occurred in the  
451 surface mixed layer. In some casts a bifurcation of the two indices was apparent (a). At higher  
452 levels of overdispersion, the Lloyd index shows a stronger numerical response than the Index of  
453 Aggregation, which means it is more sensitive to detecting patchy distributions of particles in the  
454 water column.

455

456 Fig. 7. Depth distribution of particle numbers (means per meter) versus depth (m). The inserts  
457 show the periodogram power spectral densities after fast Fourier transformation. (a) Example  
458 with no periodicity. (b) The most frequently encountered case (40 of 84casts), in which a  
459 periodicity was detectable at depth intervals shown in Fig. 8. (c) A large subsurface particle peak  
460 in some casts masked possible underlying periodicities at higher frequencies.

461

462 Fig. 8. Frequency distribution of peak periodicity of all casts ( $n=40$ ) that had a detectable  
463 frequency in the particle distribution through the water column. In one fourth of the peaks in the

464 FFT periodogram (i.e., the basic modulus-squared of the discrete Fourier transform,  $n=10$ ), the  
465 peak fell in the 30 m depth bin. This means that particle peaks occurred repeatedly at  $\sim 30$  m  
466 depth intervals in the water column.

467

468 Fig. 9. Total particle volume at two locations in the Ross Sea and at two points in time (black vs  
469 green symbols). (a) Good agreement between total particle volume and fluorescence indicating  
470 that fresh phytoplankton dominate the flux. (b) More variable relationship between total particle  
471 volume and fluorescence that also changed significantly over a period of 2 weeks.

472

473 Fig. 10. Examples of DIHM used in our analysis. (a-c) stringer-type marine snow particles, (d-e)  
474 amorphous marine snow particles held together by optically transparent exopolymers, (f-g)  
475 diatoms embedded in marine snow, (h) *Corethron* spp., (i-j) chain forming diatoms, (k)  
476 *Rhizosolenia* sp., (l-m) acantharians, (n-o) tintinnids, (p) copepod, (q) *Fritillaria* sp., (r) nauplius,  
477 (s) crustacean carcass, (t) krill fecal pellet.

478

479 Fig. 11. Pie charts of the relative contribution of various particle groups to total particle numbers  
480 below the pycnocline. We consider these particles exported as only the largest zooplankton (too  
481 rare to be accounted for by the DIHM) would be able to swim back into the upper mixed layer.

482

483 Fig. 12. Mean (numbers liter<sup>-1</sup>) (a) and integrated (m<sup>-2</sup>) (b) abundance of *Phaeocystis antarctica*  
484 colonies as determined by the DIHM. Numbers in (a) represent the depths of the mixed layer at  
485 each site.

486

487 Fig. 13. *Phaeocystis antarctica* colony abundances (numbers liter<sup>-1</sup>) along a section of the Ross  
488 Sea.

489

490 Fig. 14 Vertical distribution of *Phaeocystis antarctica* colonies (numbers liter<sup>-1</sup>) in relation to  
491 sigma<sub>θ</sub>. *P. antarctica* colonies reached much deeper when the density gradients eroded (c and d)  
492 approximately 2 weeks later. The red lines represent 15 m moving averages.

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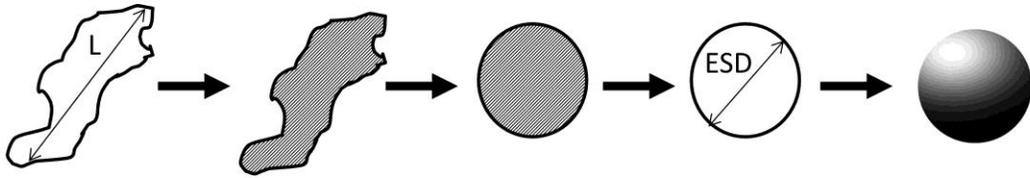


Fig. 1

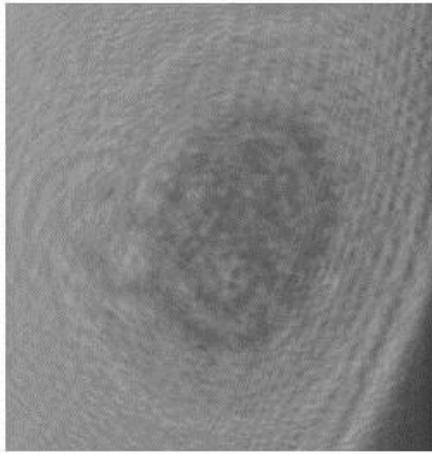


Fig. 2

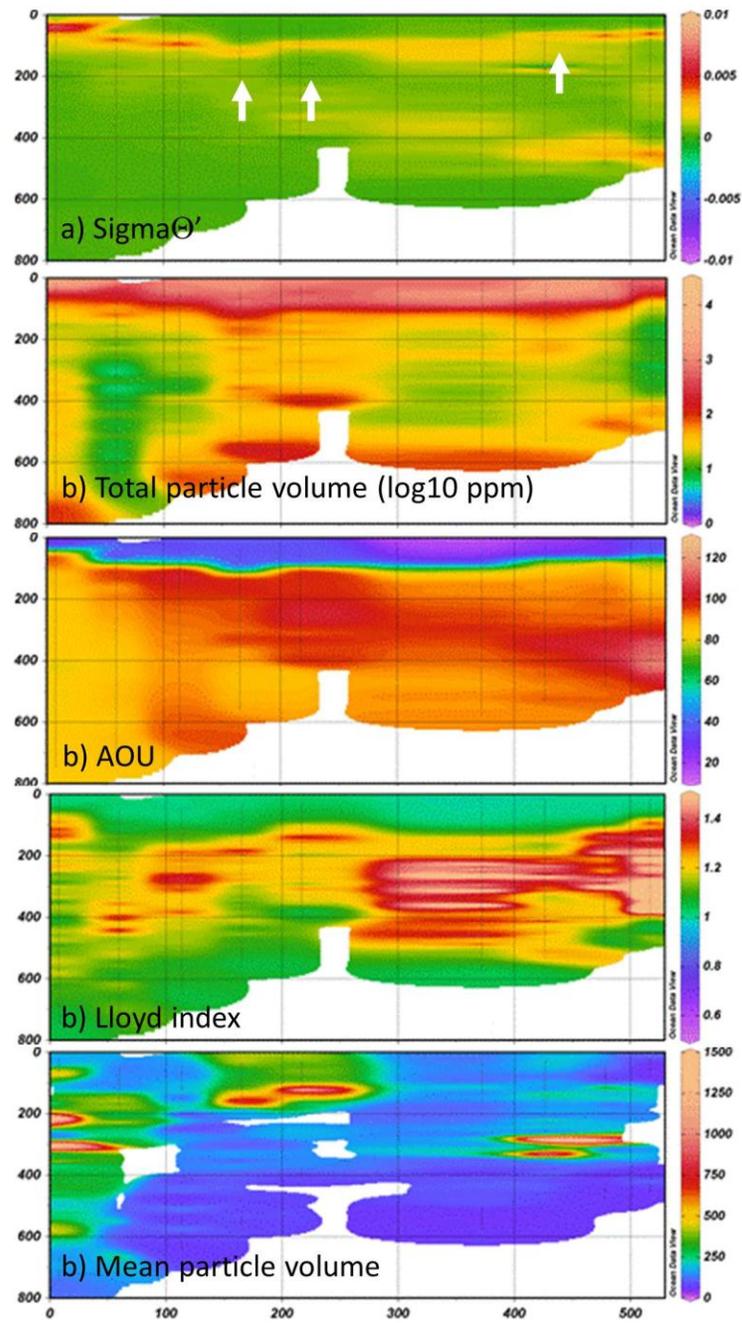


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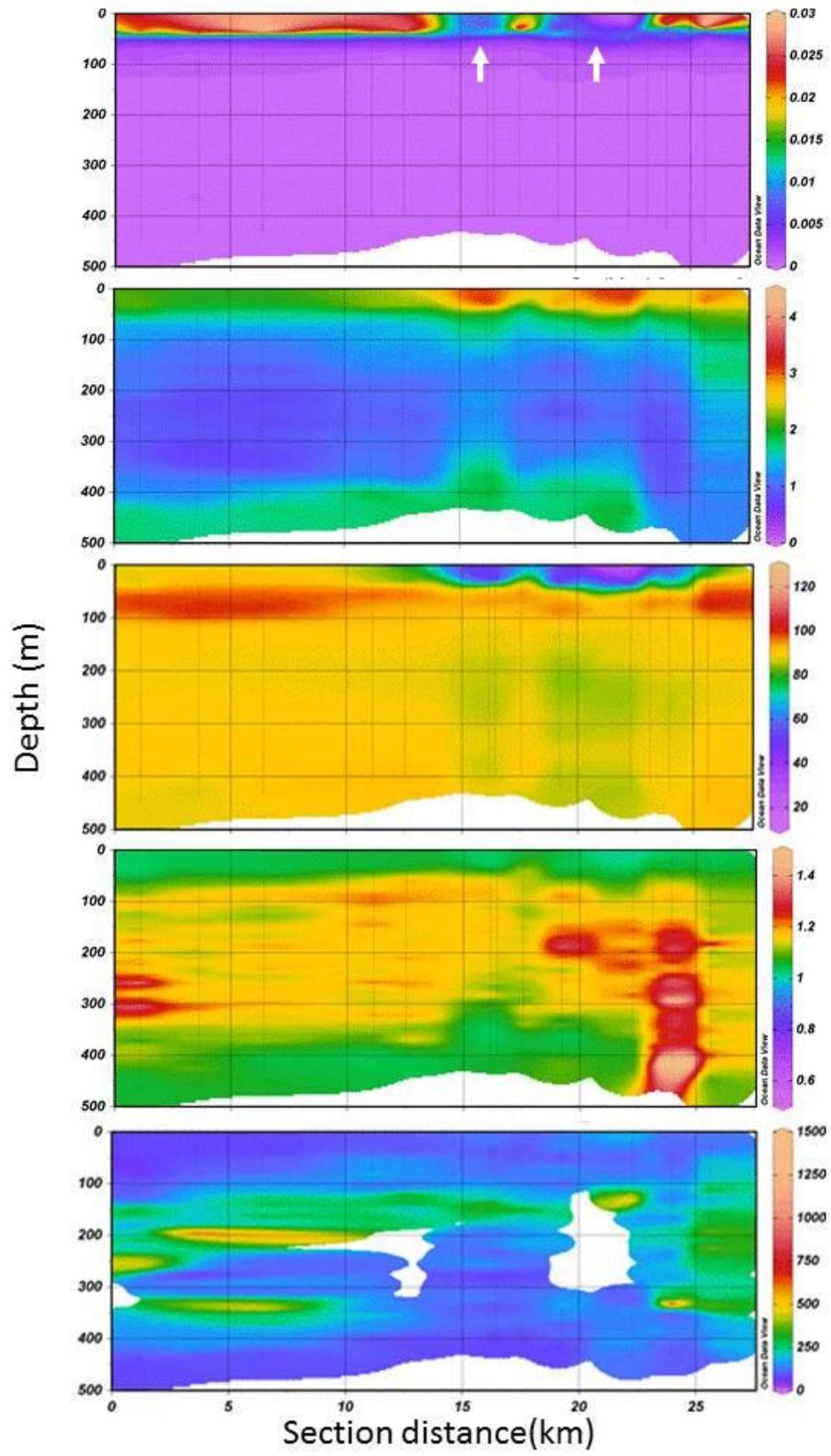


Fig. 4

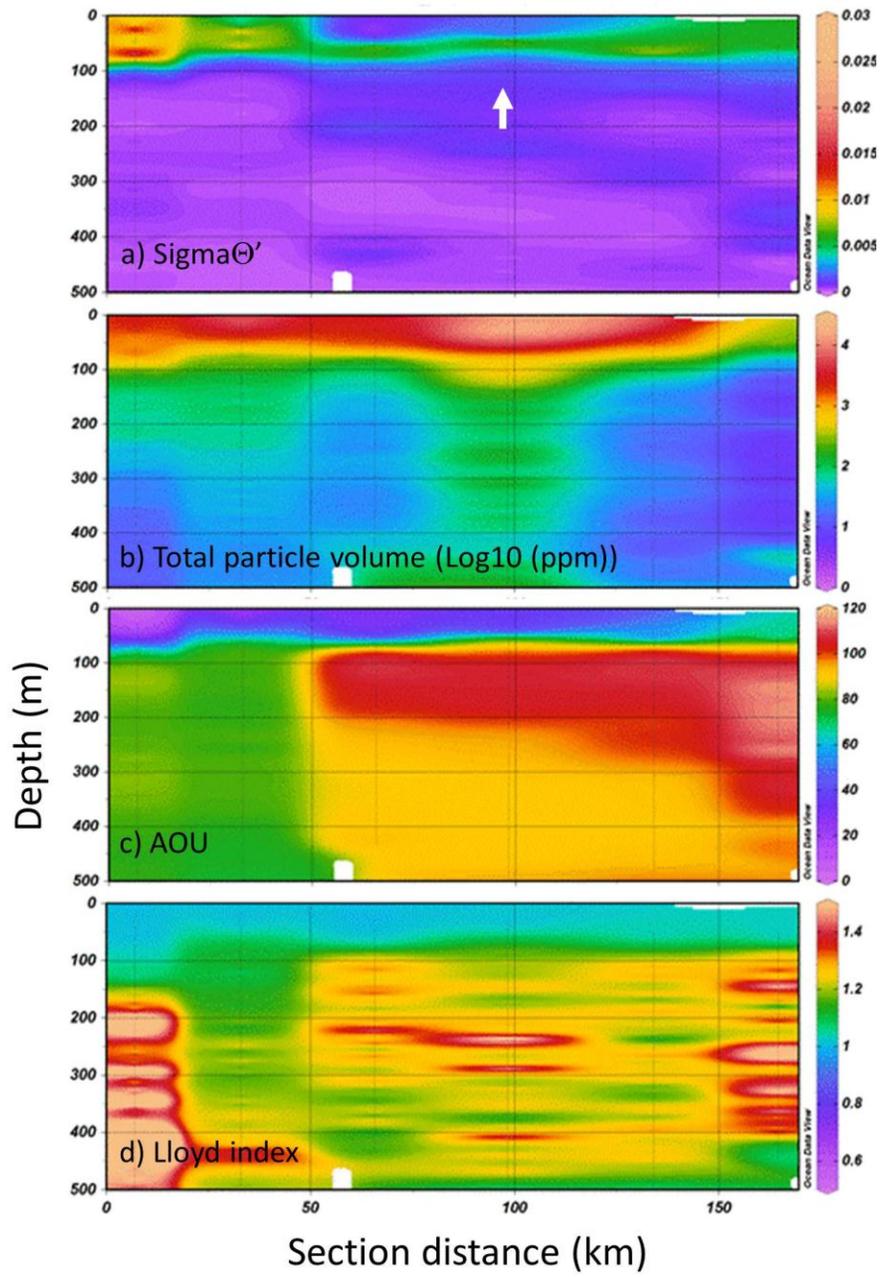


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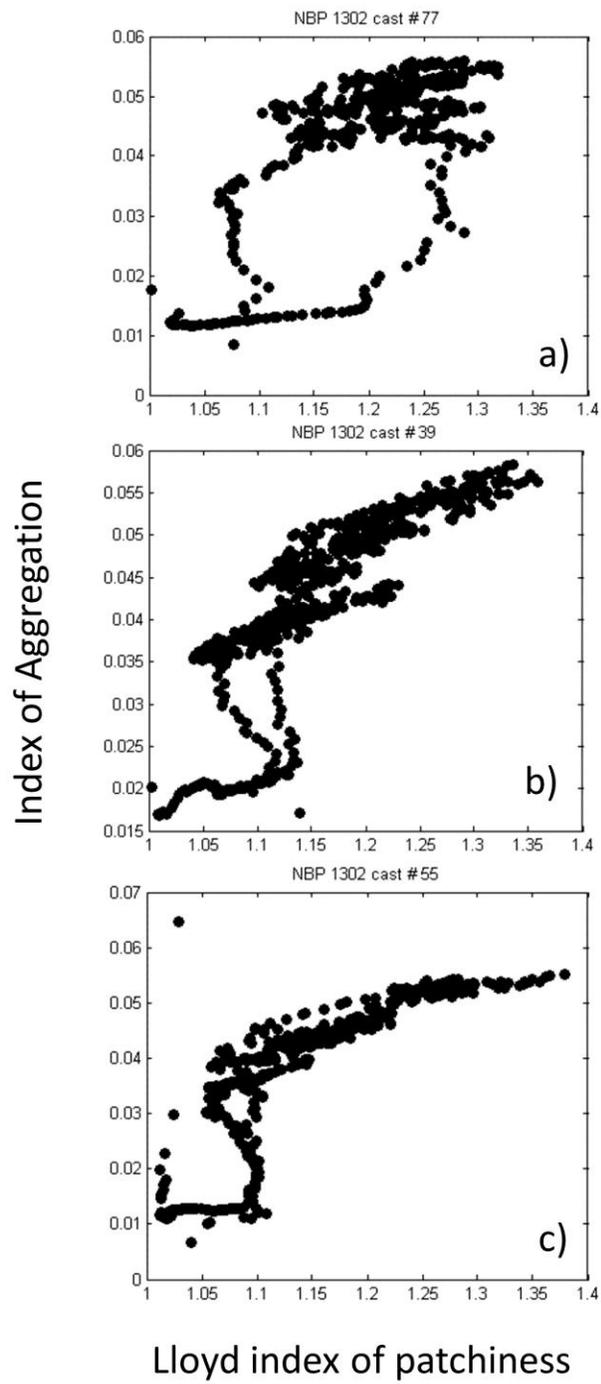


Fig. 6

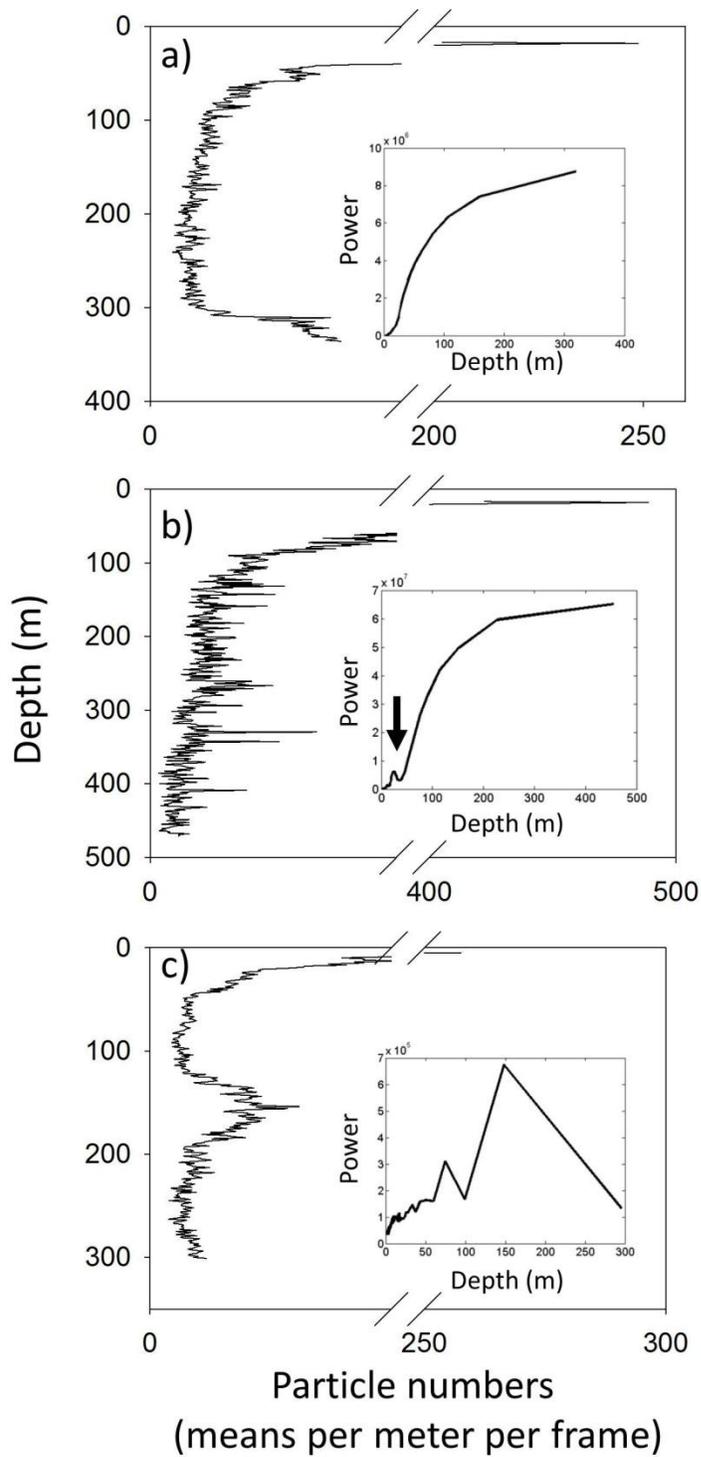


Fig. 7

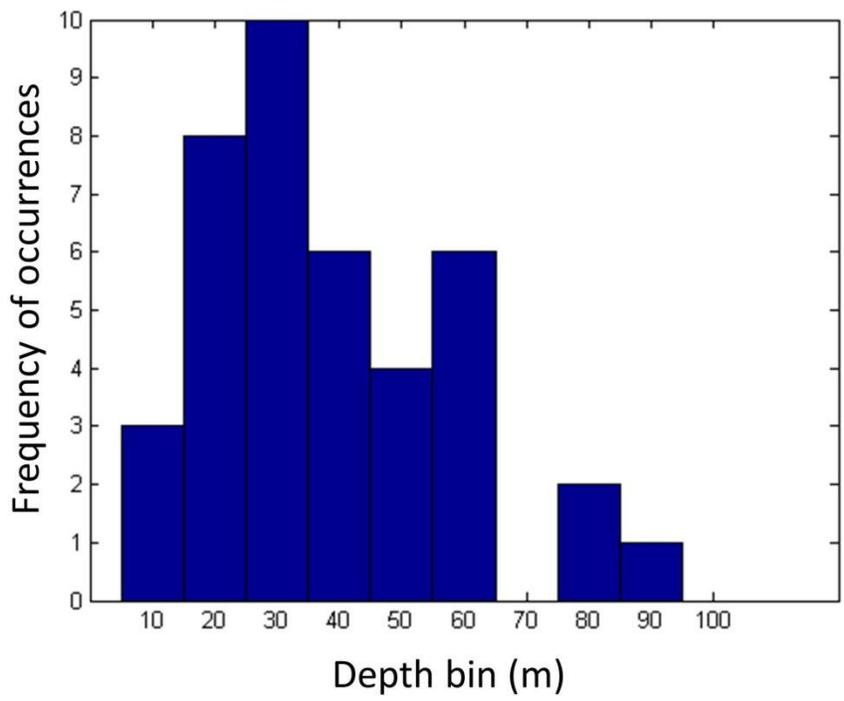


Fig. 8

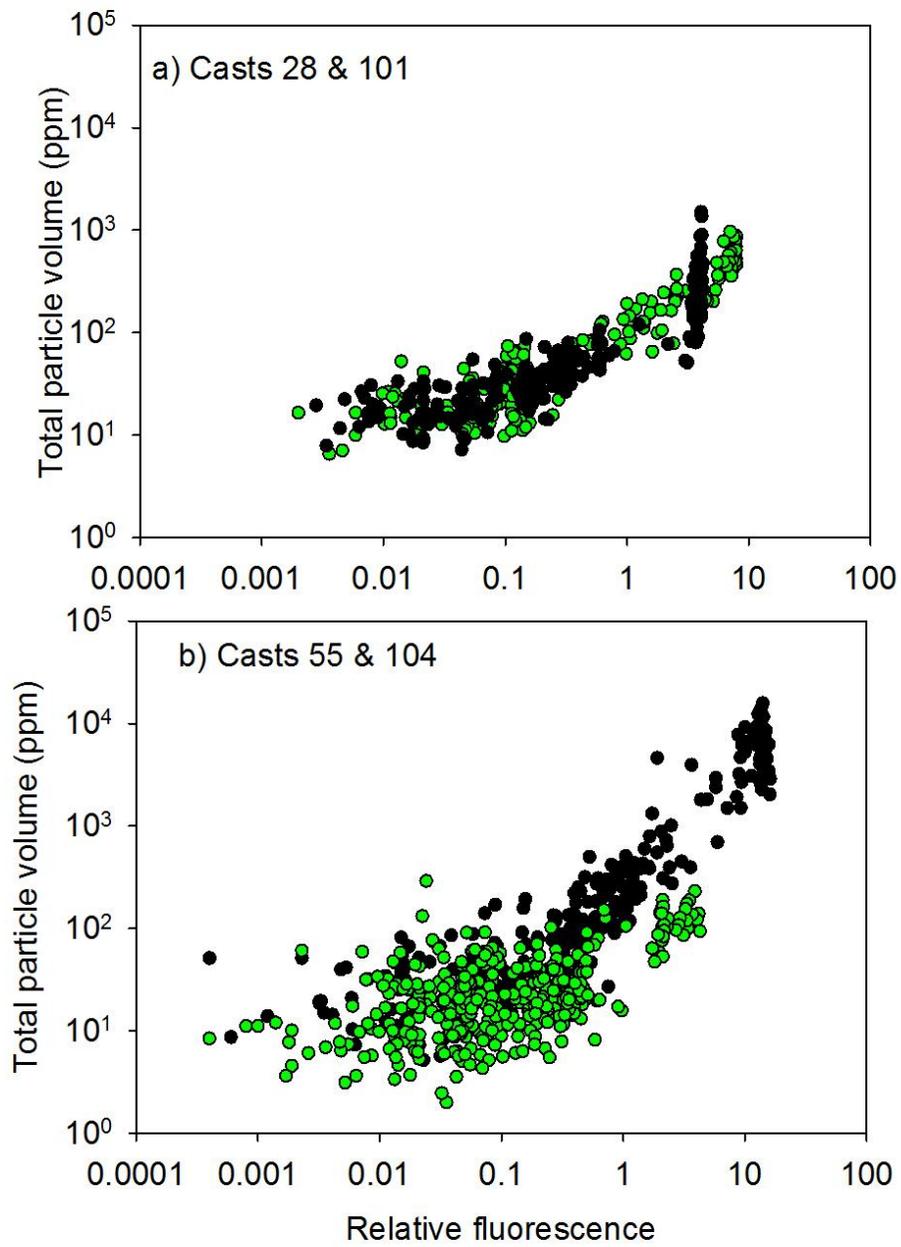


Fig. 9

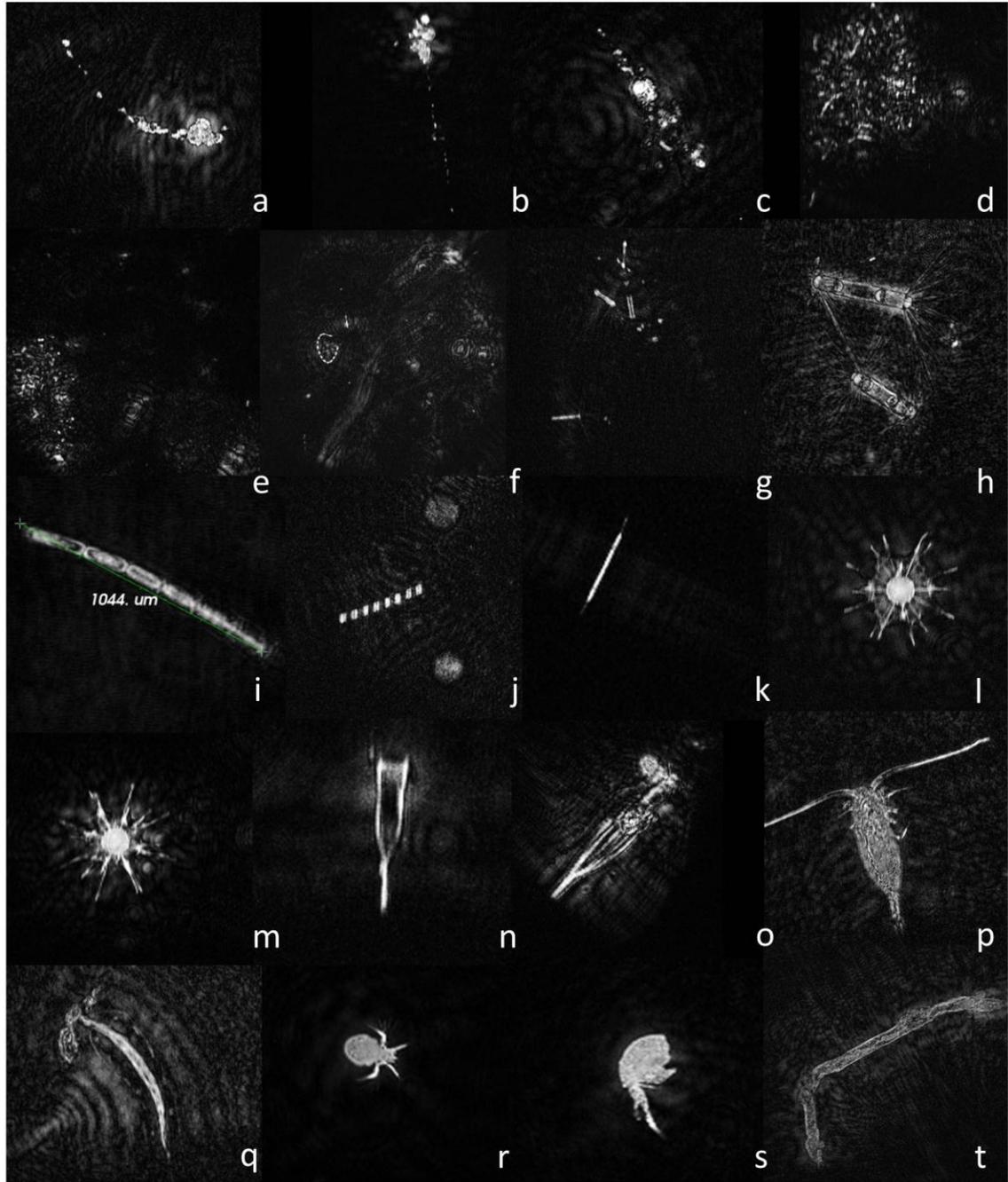


Fig. 10

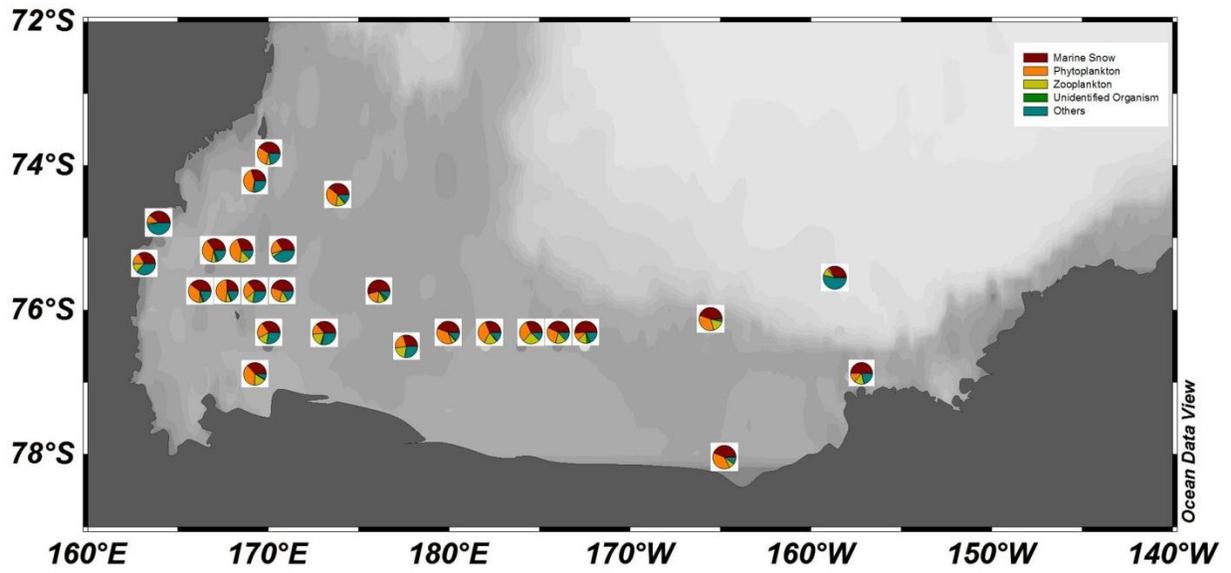


Fig. 11

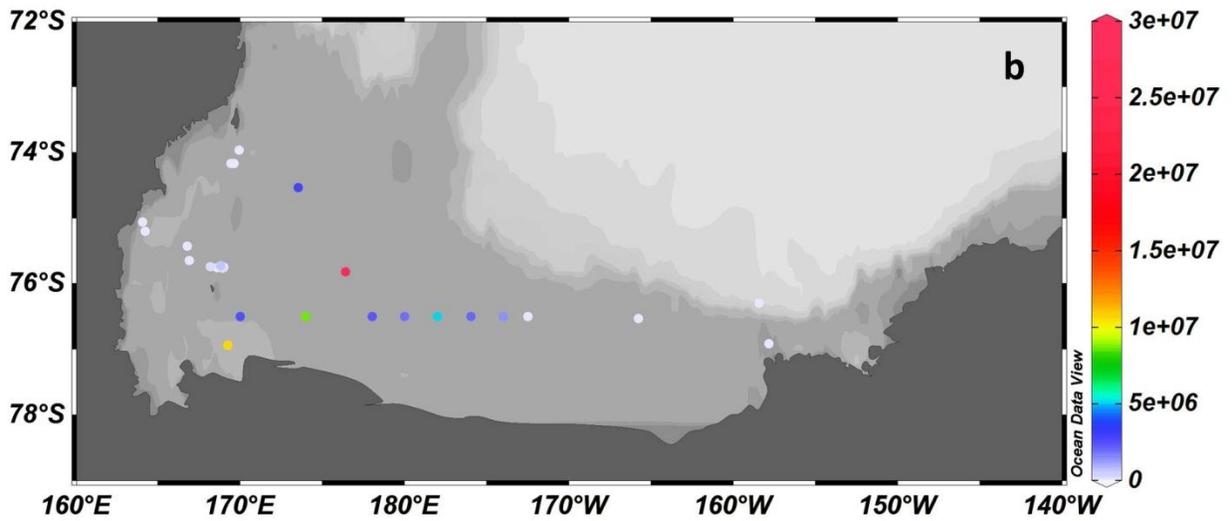
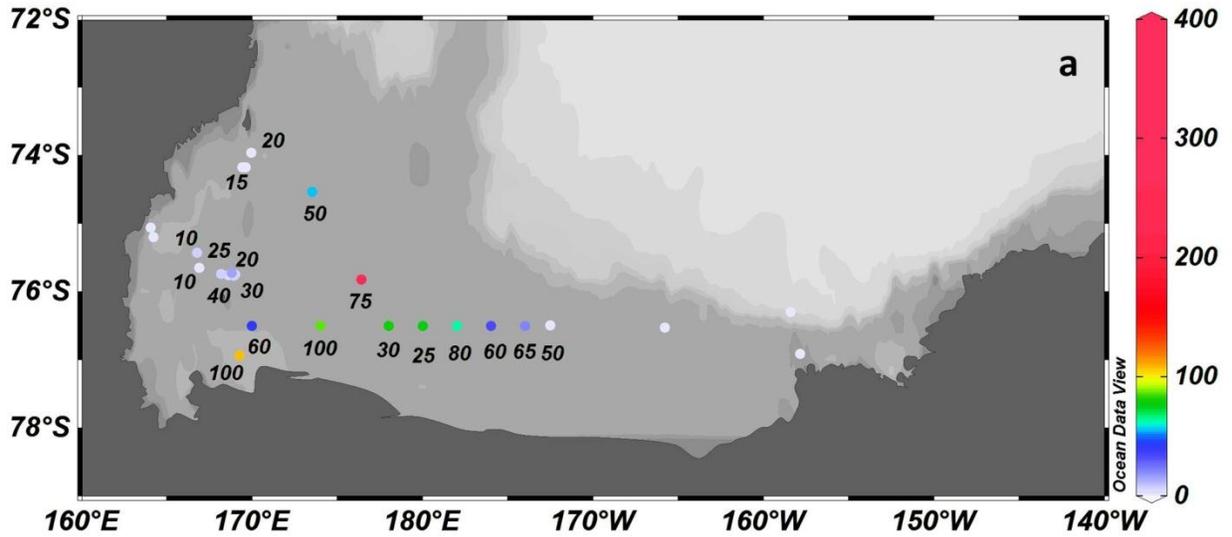


Fig. 12

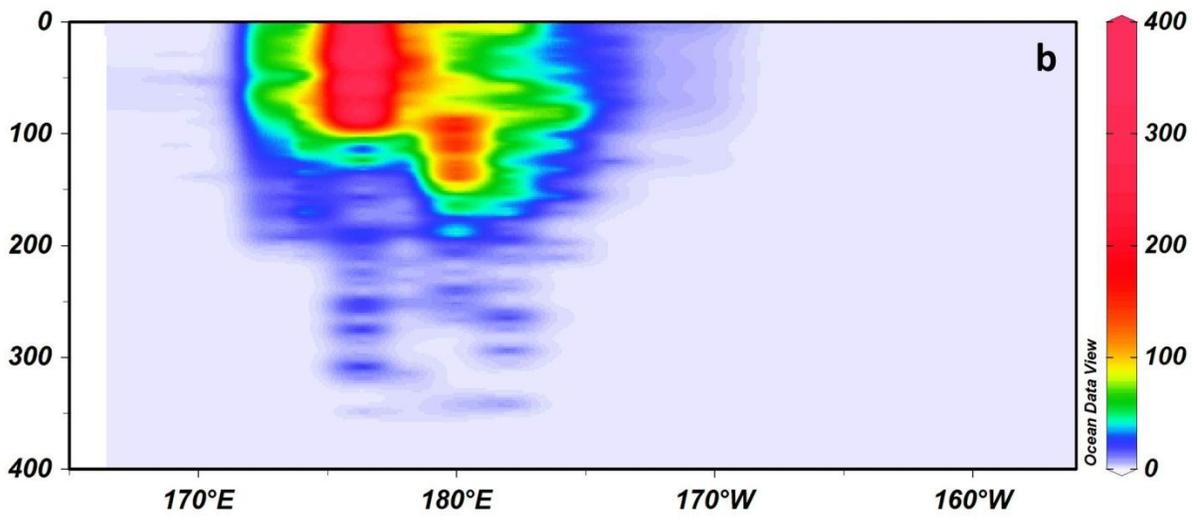
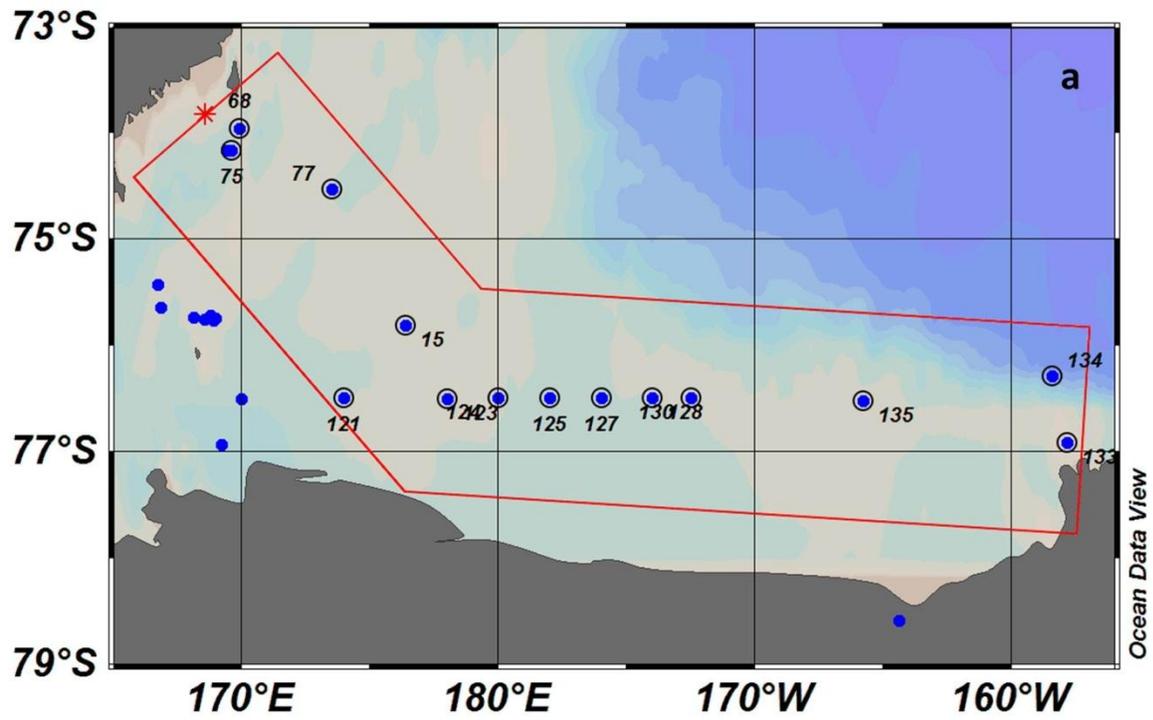


Fig. 13

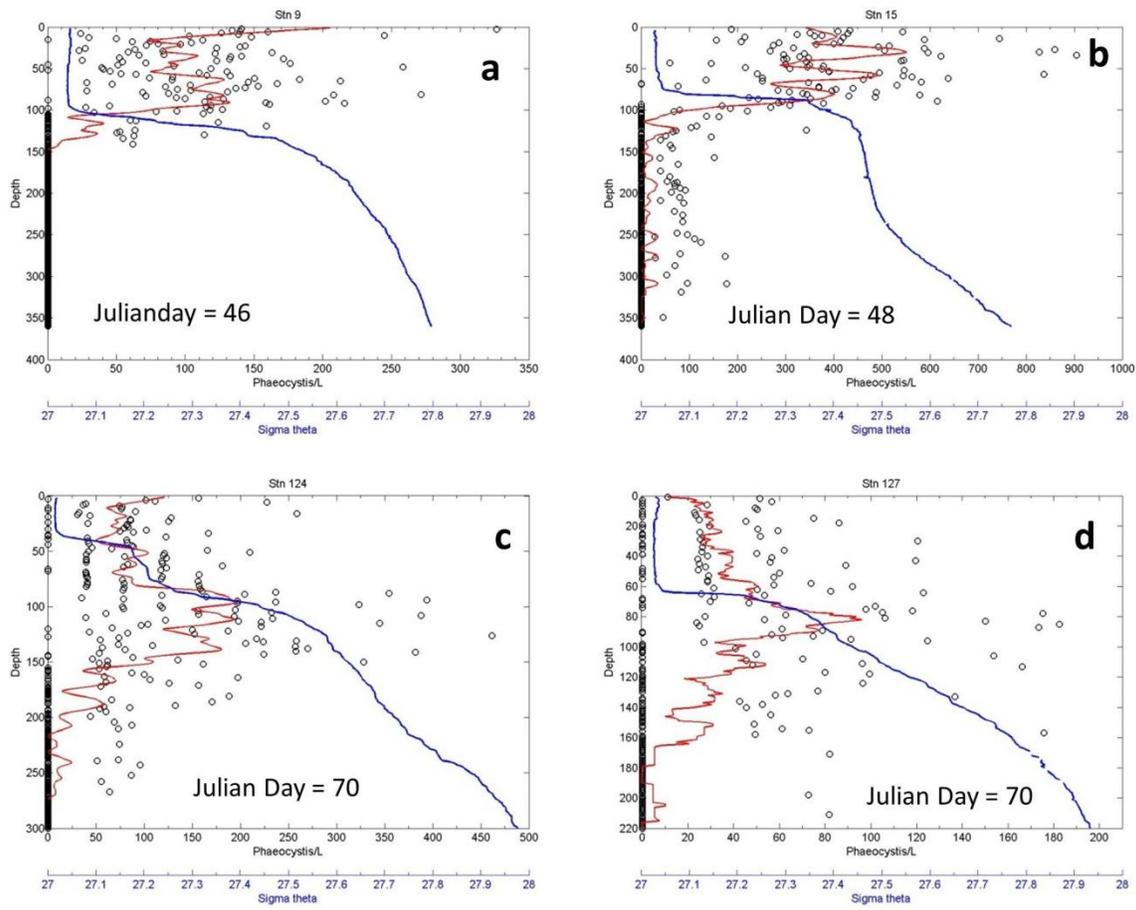


Fig. 14